Finding Holes in Multivariate Data

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The goal is to find a hole in a *k*-dimensional cloud of points. By a hole we mean a local anomaly in the data where a *k*-dimensional elliptical region within the cloud contains fewer data points than expected, given the data points in the surrounding region.

First a famous example: The Pollen Data Set. An artificial data set created by David Coleman for the 1987 JSM data competition.



Two basic ideas

- Increase the data weight in the neighborhood of the hole.
- Scatterplots that down-weight points by distance from the x-y plane.

The Pollen data again:



Outline

- Local weighting of data points near a hole: change of measure and importance sampling
- An objective function
- Optimization procedure
- Representing an *m*-dimensional hole in two dimensions
- Woods Hole tide data

Local Weighting Using Classical Multivariate Normal Results

1. Multivariate normal density

$$n(\mathbf{x};\mu,\Sigma) = \frac{1}{(2\pi)^{k/2}} \exp\left(-\frac{1}{2}(x-\mu)^{t}\Sigma^{-1}(x-\mu)\right)$$

2. Weighting function

$$g(\mathbf{x};\boldsymbol{\mu}_{w},\boldsymbol{\Sigma}_{w}) = \exp\left(-\frac{1}{2}\left(x-\boldsymbol{\mu}_{w}\right)^{t}\boldsymbol{\Sigma}_{w}^{-1}\left(x-\boldsymbol{\mu}_{w}\right)\right)$$

Change of Measure Importance Sampling

Standard Bayes calculation for the multivariate normal

Local normal density

$$n(\mathbf{x};\boldsymbol{\mu}_l,\boldsymbol{\Sigma}_l) = C \quad n(\mathbf{x};\boldsymbol{\mu},\boldsymbol{\Sigma})g(\mathbf{x};\boldsymbol{\mu}_w,\boldsymbol{\Sigma}_w)$$

Local Covariance Matrix

$$\boldsymbol{\Sigma}_l = \left(\boldsymbol{\Sigma}^{-1} + \boldsymbol{\Sigma}_w^{-1}\right)^{-1}$$

Local Mean

$$\boldsymbol{\mu}_{l} = \boldsymbol{\Sigma}_{l} \left(\boldsymbol{\Sigma}^{-1} \boldsymbol{\mu} + \boldsymbol{\Sigma}_{w}^{-1} \boldsymbol{\mu}_{w} \right)$$

Normalizing constant

$$C^{-1} = \frac{\left|\Sigma_{l}\right|^{\frac{1}{2}}}{\left|\Sigma\right|^{\frac{1}{2}}} \exp\left(+\frac{1}{2}\left(\mu_{l}^{t}\Sigma_{l}^{-1}\mu_{l}-\mu^{t}\Sigma^{-1}\mu_{l}-\mu_{w}^{t}\Sigma_{w}^{-1}\mu_{w}\right)\right)$$



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General Hole Finding Procedure

1. For fixed weight parameters μ_w and Σ_w compute the weights

$$w_i = \frac{g(x_i, \boldsymbol{\mu}_w, \boldsymbol{\Sigma}_w)}{\sum_{i=1}^n g(x_i, \boldsymbol{\mu}_w, \boldsymbol{\Sigma}_w)}$$

2. Estimate the local parameters

$$\hat{\mu}_L = \sum_{i=1}^n w_i \mathbf{x}_i$$
 and $\hat{\Sigma}_l = \sum_{i=1}^n w_i (\mathbf{x}_i - \hat{\mu}_l) (\mathbf{x}_i - \hat{\mu}_l)^i$

3. Compute the objective function.

$$\overline{p}(\boldsymbol{\mu}_{w},\boldsymbol{\Sigma}_{w}) = \sum_{i=1}^{n} g(\mathbf{x}_{i}; \hat{\boldsymbol{\mu}}_{l}, \boldsymbol{\alpha} \, \hat{\boldsymbol{\Sigma}}_{l}) \, w_{i}$$

4. Minimize $\overline{p}(\mu_w, \Sigma_w)$ with respect to μ_w and Σ_w .

Numerical Details

1) Parameters are constrained so that $\sum_{i=1}^{n} g(x_i, \mu_w, \Sigma_w) = n_w << n$

In these examples n_{w} is less than 10% of the data.

2) The Cholesky decomposition is used to parameterize the matrix

$$\Sigma_w^{-1} = LL^t$$

3) We use R, nlminb and selected starting values to search for the minimum.

Scatterplot Displays for *k*-dimensional Holes

- 2-dimensional holes in *k*-dimensional space is just a projection-pursuit problem.
- Random points on a *k*-dimensional sphere.
- Scatterplots for for *k*-dimensional holes.



Model Hole in X1 X2 Plane: Hole labeled by distance



Orthogonal Transform of Model Data : Hole labeled by distance

Classic Example, Random points on the unit Sphere

Sphere Dimension 3

Sphere Dimension 5



Solution: Distance from the x-y plane

$$d_i = \sum_{j=3}^k x_{ij}^2 .$$

Down-weight points by

$$\eta_i = \exp(-c\,d_i),$$

where c denotes a constant dependent on dimension and point density.



Down Weighted by Distance

Sphere the Locally Weighted Data

Weighted mean and covariance

$$\widehat{\boldsymbol{\mu}}_{l} = \sum_{i=1}^{n} w_{i} \mathbf{x}_{i} \quad and \quad \widehat{\boldsymbol{\Sigma}}_{l} = \sum_{i=1}^{n} w_{i} (\mathbf{x}_{i} - \widehat{\boldsymbol{\mu}}_{l}) (\mathbf{x}_{i} - \widehat{\boldsymbol{\mu}}_{l})^{t}$$

Sphered data, centered at 0 and with unit sample covariance matrix $\mathbf{x}_{i}^{*} = (x_{i} - \hat{\mu}_{l})V_{l}\Lambda^{-1/2}$,

 V_l denotes the matrix of eigen vectors of $\hat{\Sigma}_l$.

 Λ denotes the diagonal matrix of eigen values of $\hat{\Sigma}_l$.

Final step: Standard Projection Pursuit

Find a *k*-by-2 orthonormal matrix *M* that minimizes the weighted projected points in the center of the (y_1, y_2) plane. project $y_i = x_i^* M$ minimize $\sum_{i=1}^n w_i \exp(-ay_i y_i^t)$ Woods Hole Tide Gauge data

- 1. Woods Hole hourly water level measurements.
- 2. Measurements lagged by 1 through 5 hours.
- 3. The data was thinned by using every third observation.



Hurricane Sandy Data NOAA web





Woods Hole Tide Data: First Three Principal Components



Hole Models are essentially random truncation models. For one-dimensional holes see Morrell, C. H. and Johnson R. A.(1991), "Random truncation and neutrinos", Technometrics:,33, 429-440.

Pollen Data

Becker, R. A., L. Denby, R. McGill and A. R. Wilks(1986). Datacryptanalysis: A Case Study. *Proceedings of the Section on Statistical Graphics of Amer. Statistical Association*, 1987, pp. 92-97. (And a Bell Labs Tech Report)

Found the three holes. But only because of a special property of the simulation.In the simulation all data points were paired.Truncated points were then the missing points in a pair.The missing points formed three clusters.

Finding Holes in Data using two- and three-dimensional non parametric density estimation.

Scott, D.W. (2009). Multivariate Density Estimation: Theory, Practice, and Visualization, John Wiley, New York.

Nested α -level density contours indicates a hole.

The general consensus seems to be that finding holes in data is

Too Hard and Too Esoteric. Locally-Weighted Hole Finder

- Is robust to departures from multivariate normality.
- Is based on classical multivariate analysis.
- Solves a rather hard problem in a transparent and straightforward way.
- Generalizable to other departures from "Local Normality".

These slides are on my web site at Temple University

astro.temple.edu/wk**smith**/